

Hot Jupiters & Cool Stars

Beware:

A new TiO line list is here.

Laura K. McKemmish

University of New South Wales

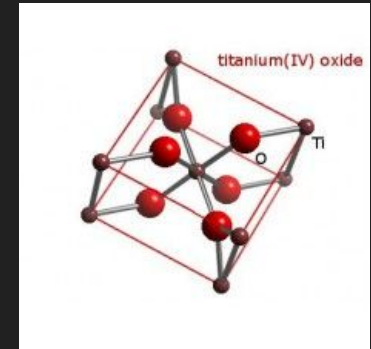
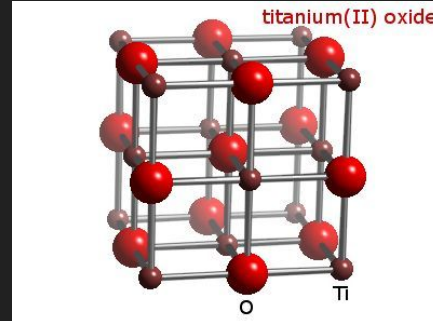
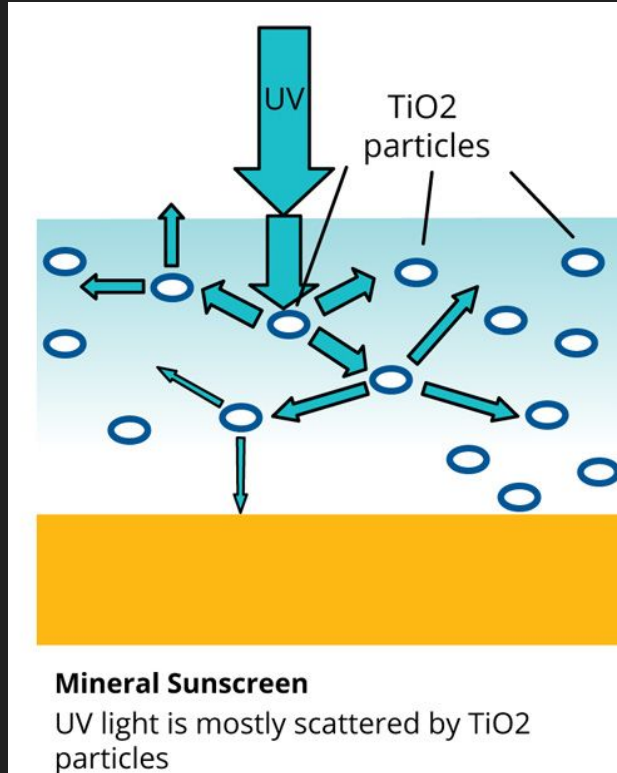
Thomas Masseron

Instituto de Astrofísica de Canarias

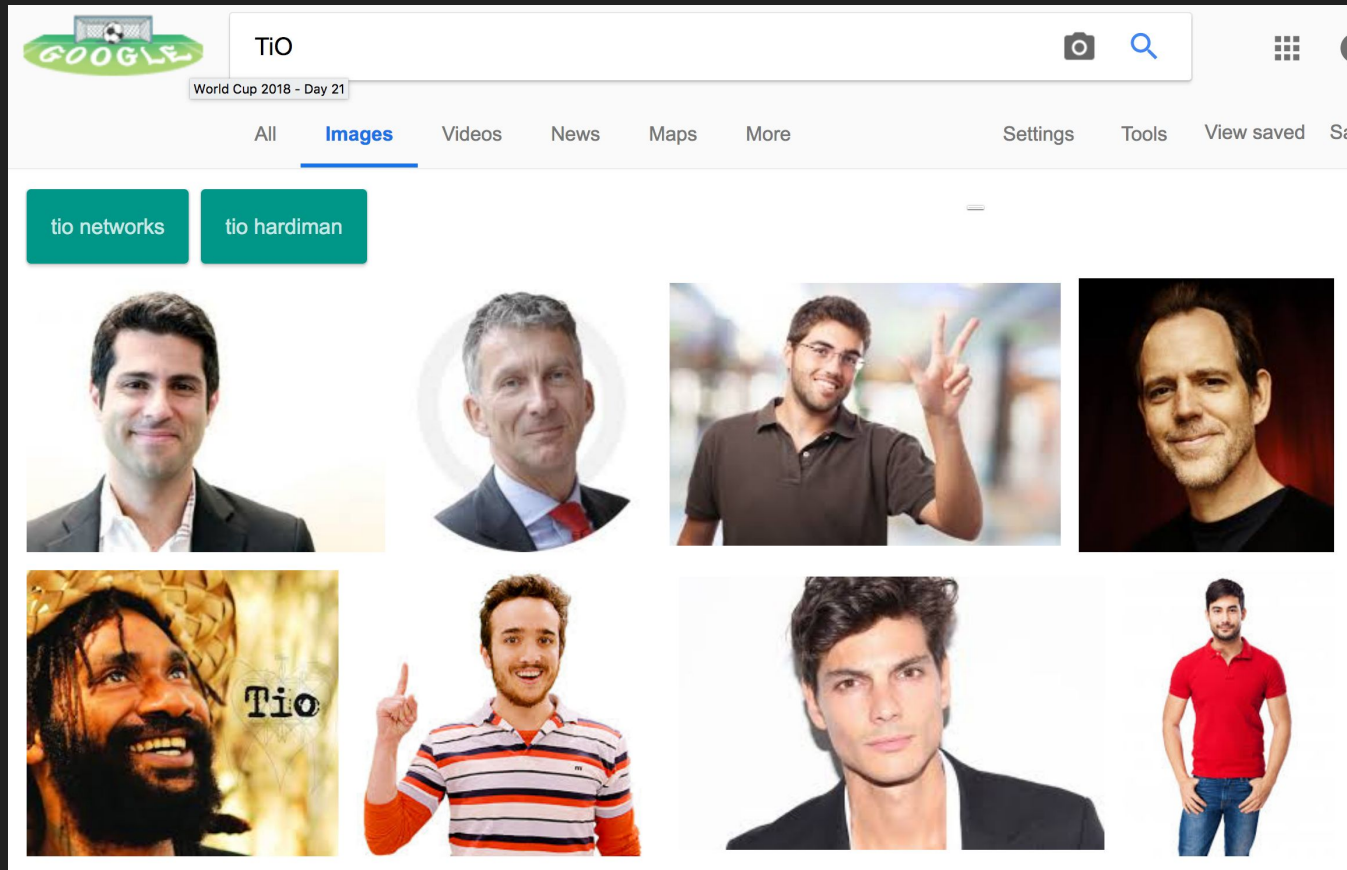
Sergey N. Yurchenko, Jonathan Tennyson

University College London

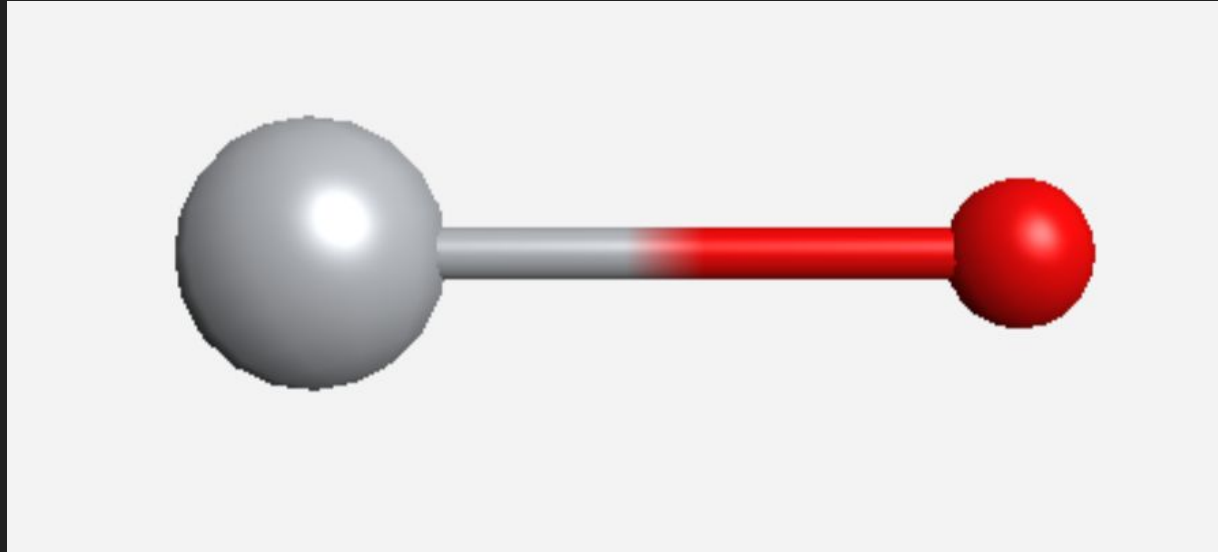
Titanium oxide to chemists



TiO to Google

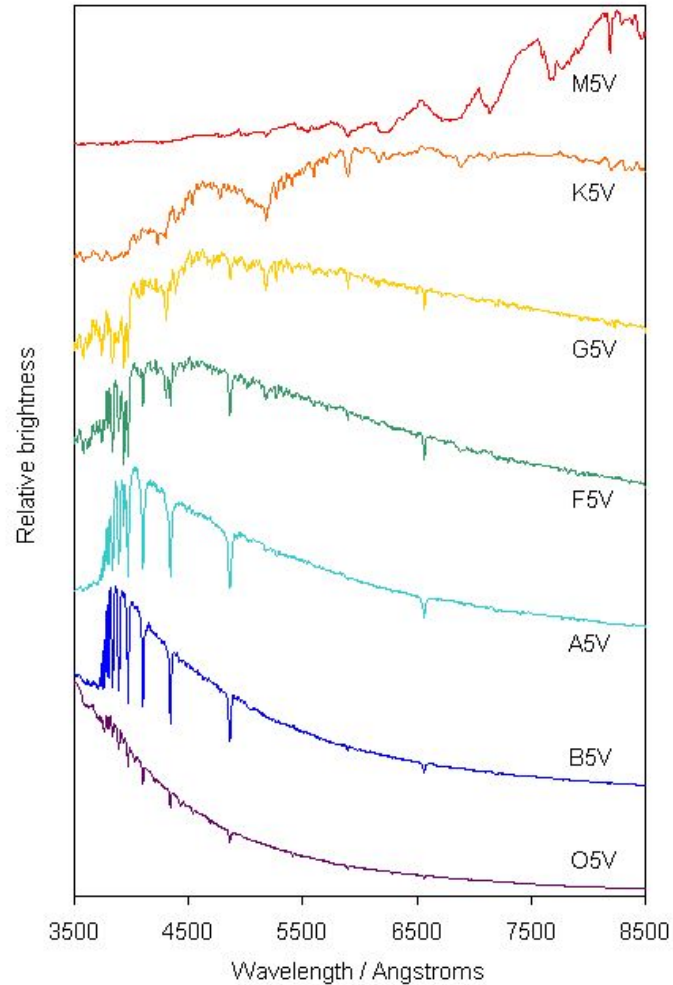


Titanium oxide to astronomers

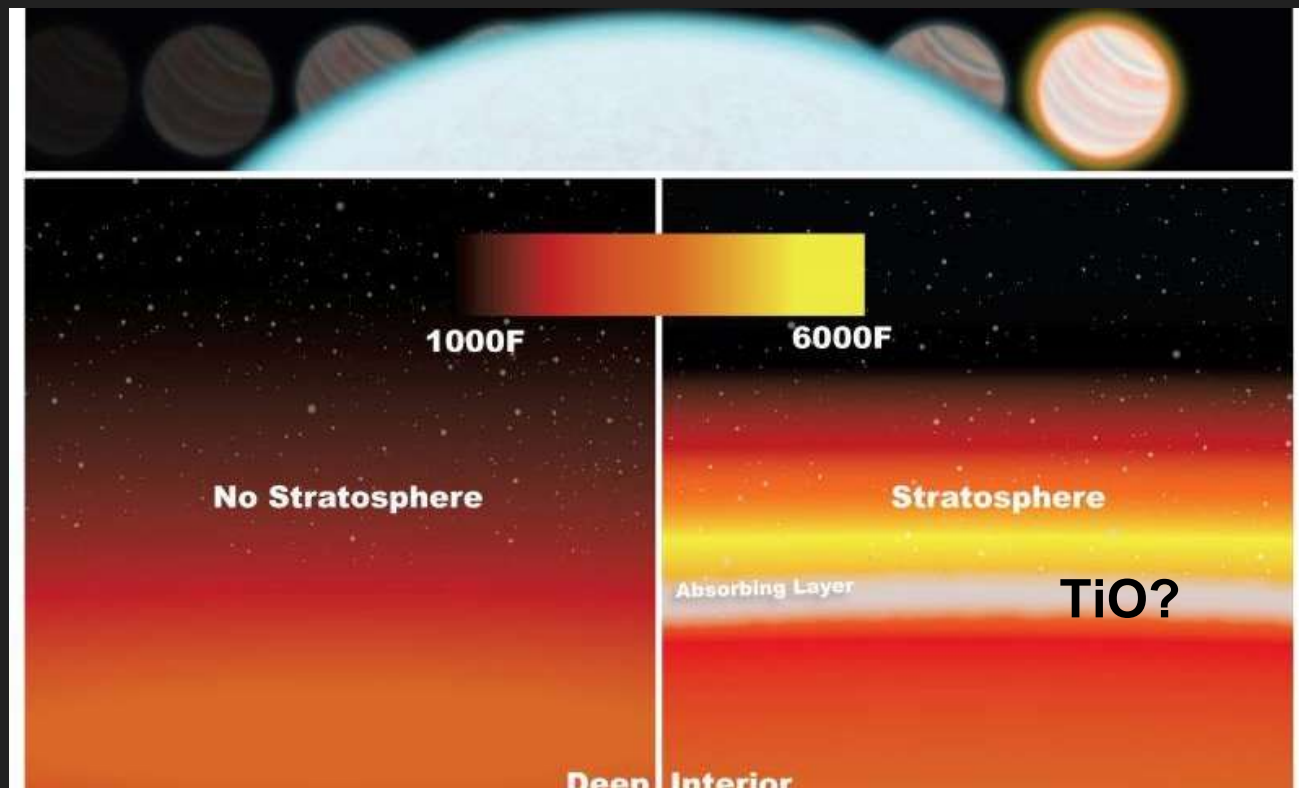


One of top 10 most important molecules

TiO in stars



TiO, hot Jupiters and temperature inversion



Hot Jupiter: Recent Results

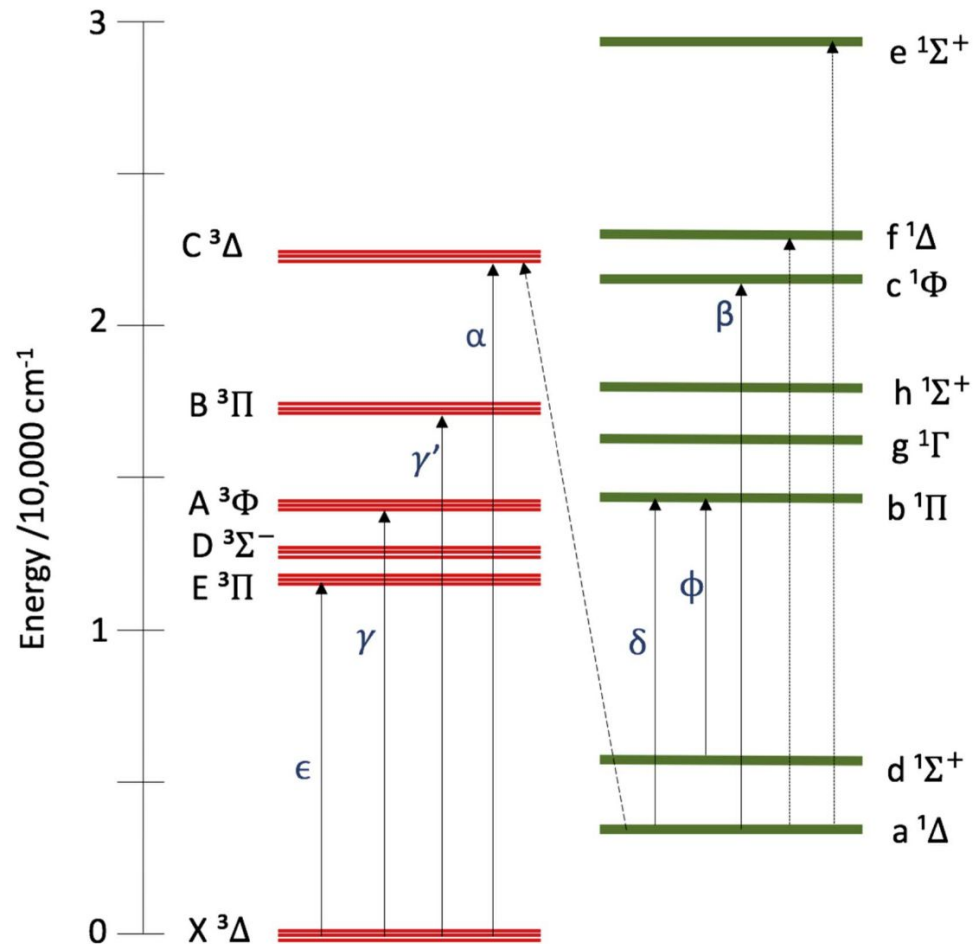
Letter | Published: 13 September 2017

Detection of titanium oxide in the atmosphere of a hot Jupiter

Elyar Sedaghati , Henri M. J. Boffin, Ryan J. MacDonald, Siddharth Gandhi, Nikku Madhusudhan, Neale P. Gibson, Mahmoudreza Oshagh, Antonio Claret & Heike Rauer

Nature **549**, 238–241 (14 September 2017) | [Download Citation](#) 

Electronic structure of TiO



Schwenke 1998

Faraday Discuss., 1998, **109**, 321–334

Opacity of TiO from a coupled electronic state calculation parametrized by *ab initio* and experimental data

David W. Schwenke

M.S. 230-3, NASA Ames Research Center, Moffett Field, CA 94035-1000, USA

- 13 coupled electronic states considered

Plez 1998

Astron. Astrophys. 337, 495–500 (1998)

ASTRONOMY
AND
ASTROPHYSICS

A new TiO line list

B. Plez^{1,2}

¹ Astronomiska Observatoriet, Box 515, S-751 20 Uppsala, Sweden

² Atomspektroskopi, Fysiska Institution, Box 118, S-221 00 Lund, Sweden (plez@fysik.lu.se)

- 9 uncoupled electronic states considered

Plez 1998 ... improved in VALD (2012)

Astron. Astrophys. 337, 495–500 (1998)

ASTRONOMY
AND
ASTROPHYSICS

A new TiO line list

B. Plez^{1,2}

¹ Astronomiska Observatoriet, Box 515, S-751 20 Uppsala, Sweden

² Atomspektroskopi, Fysiska Institution, Box 118, S-221 00 Lund, Sweden (plez@fysik.lu.se)

- 9 uncoupled electronic states considered
- **Incorporates experimental frequencies where available**
 - Shifted isotopologue frequencies by error in ^{48}TiO

High Resolution Ground-based Exoplanet Spectroscopy



Call to action

A&A 575, A20 (2015)

A search for TiO in the optical high-resolution transmission spectrum of HD 209458b: Hindrance due to inaccuracies in the line database

H. J. Hoeijmakers¹, R. J. de Kok^{1,2}, I. A. G. Snellen¹, M. Brogi^{1,*}, J. L. Birkby^{1,3,} and H. Schwarz¹**

Resolutions of 100,000

At $1 \text{ } \mu\text{m} = 1000 \text{ } \text{\AA} = 10,000 \text{ cm}^{-1}$, this is equivalent to $0.01 \text{ } \text{\AA} = 0.1 \text{ cm}^{-1}$



Creating a New TiO Linelist

Methods and Tools

Full
Spectroscopic
Model

Solve for motion of the **electrons**

Molpro

Energy Spectroscopic Model

Potential
energy
curves

Spin-orbit &
other
couplings

Dipole
moment
curves

Solve for motion of the **nuclei**

Duo

Rovibronic
Wavefunctions

Rovibronic Energies

Transition Intensities

ExoMol

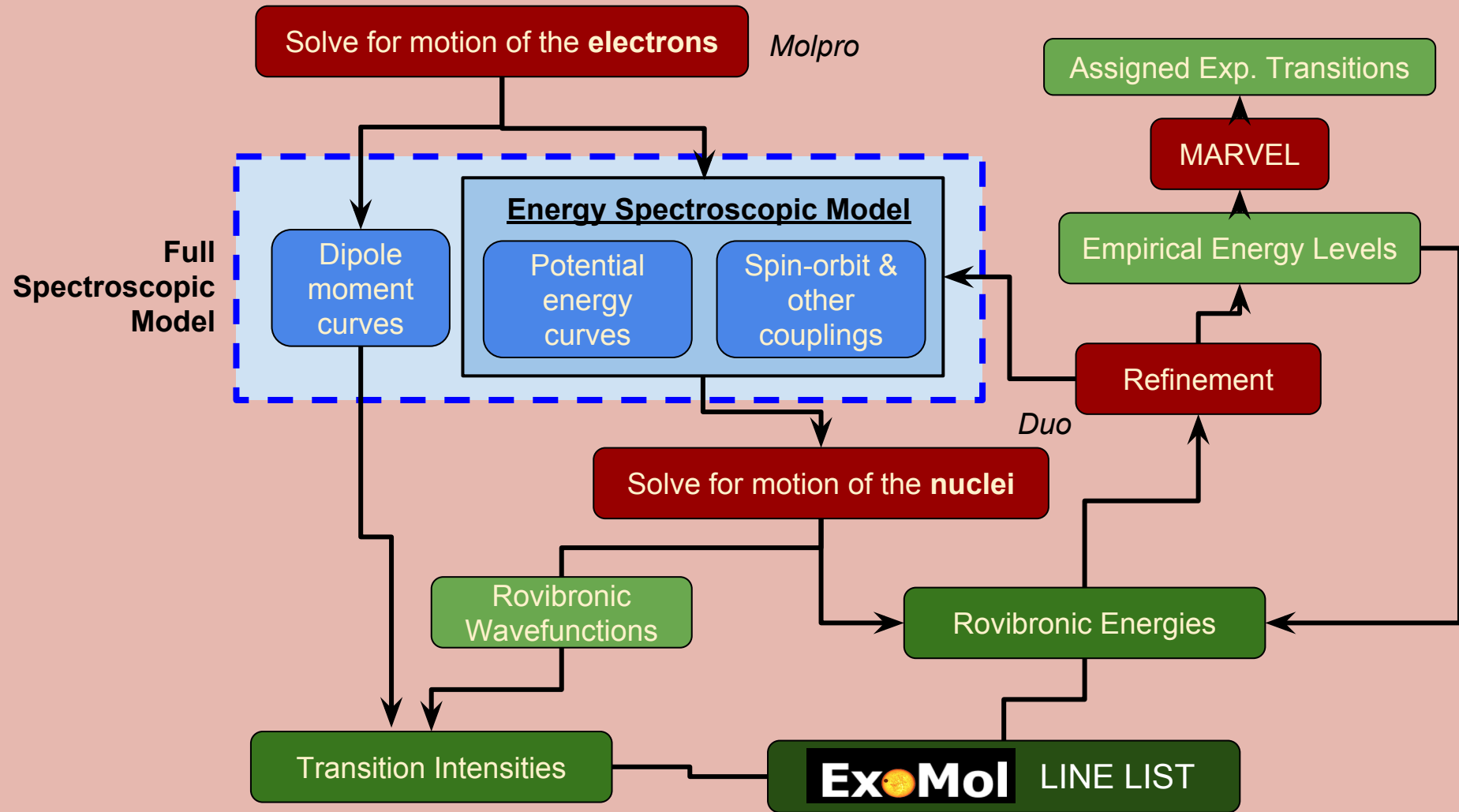
LINE LIST

Assigned Exp. Transitions

MARVEL

Empirical Energy Levels

Refinement



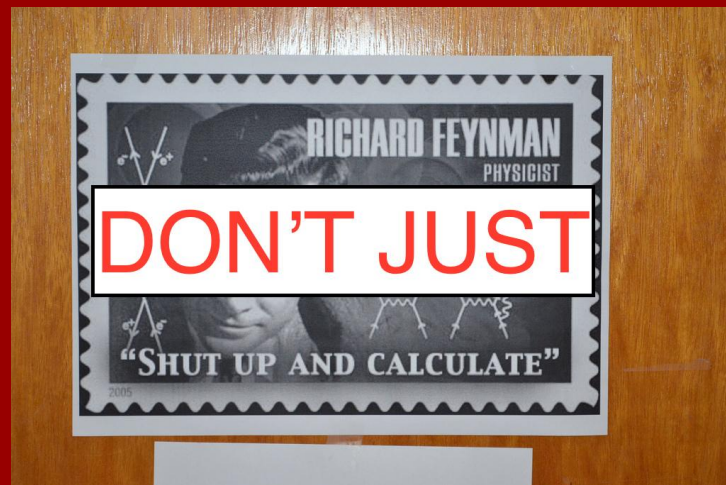
Ab-initio (Molpro)

Solve for motion of the **electrons**

Dipole
moment
curves

Potential
energy
curves

Spin-orbit &
other
couplings



TOPICAL REVIEW

The *ab initio* calculation of spectra of open shell diatomic molecules

Jonathan Tennyson , Lorenzo Lodi, Laura K McKemmish and Sergei N Yurchenko

Published 5 May 2016 • © 2016 IOP Publishing Ltd

[Journal of Physics B: Atomic, Molecular and Optical Physics](#), [Volume 49](#), [Number 10](#)

It's hard and accuracy is low

MARVEL

Assigned Exp. Transitions

MARVEL

Empirical Energy Levels



Available online at www.sciencedirect.com



Journal of Molecular Spectroscopy 245 (2007) 115–125

Journal of
MOLECULAR
SPECTROSCOPY

www.elsevier.com/locate/jms

MARVEL: measured active rotational–vibrational energy levels

Tibor Furtenbacher ^a, Attila G. Császár ^{a,*}, Jonathan Tennyson ^b



ONE PAPER & A MOLECULE

Input

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CITATIONS & GOOGLE!

Output

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MARVEL Project

ASSIGNED TRANSITIONS

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EMPIRICAL ENERGY LEVELS



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RESEARCH: DIGITALISATION

THE ASTROPHYSICAL JOURNAL SUPPLEMENT SERIES No. 247, 27:319-330
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THE FUNDAMENTAL ROTATION-VIBRATION BAND OF TiO

JOHN G. PHILLIPS
Astronomy Department, University of California, Berkeley
Received 1973 August 21

TABLE 3

$v = 0 \rightarrow v = 1$

J	P_1 cm ⁻¹	R_1 cm ⁻¹	P_2 cm ⁻¹	R_2 cm ⁻¹	P_3 cm ⁻¹	R_3 cm ⁻¹
2	996.685	1002.945				
3	996.685	1002.946*		1004.180		
4	995.640	1004.990	995.679*	1005.237*		
5	994.514*	1006.064*	994.564*	1006.268*	994.575*	1006.290
6	993.510	1007.079*	993.486*	1007.294*	993.470	1007.410*
7	992.320	1008.090*	992.381*	1008.313*	992.360	1008.439*
8	991.289*	1009.090*	991.271*	1009.327*	991.228*	1009.475
9	990.186*	1010.175	990.160	1010.410	990.135	1010.477*
10	989.120	1011.145	989.180	1011.375	988.970	1011.490
11	988.012*	1012.100	987.940	1012.390	987.845	1012.550
12	986.930	1013.140	986.820	1013.450	986.795	1013.487*
13	985.775	1014.150	985.715	1014.240	985.565	1014.495
14	984.710	1015.115	984.555	1015.315	984.385	1015.450
15	983.575	1016.125	983.350	1016.275	983.195	1016.405
16	982.405	1017.090*	982.190	1017.295	982.020	1017.370
17	981.200	1018.030	981.040	1018.210	980.845	1018.335
18	980.100	1018.910	979.905	1019.120	979.700	1019.405
19	978.895	1019.840	978.760	1020.090	978.500	1020.290
20	977.765	1020.745	977.620	1021.030	977.285	1021.200
21	976.615	1021.760	976.400	1021.985	976.060	1021.175
22	975.450	1022.665	975.125	1022.860	974.875	1022.130
23	974.245	1023.565	973.945	1023.825	973.625	1023.000
24	973.015	1024.480	972.630	1024.730	972.445	1023.975
25	971.945	1025.390	971.305	1025.640	971.235	1024.875
26	970.645	1026.285	970.305	1026.550	970.055	1025.775
27	969.610	1027.115	969.035	1027.405	968.805	1026.675
28	968.205	1028.010	967.845	1028.295	967.550	1027.545
29	967.080	1028.835	966.450	1029.185	966.305	1028.455
30	965.745	1029.755	965.410	1030.040	965.035	1029.315
31	964.555				964.195	1030.070

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	A	B	C	D	E	F	G	H
J2	τ_1 -1 cm	τ_2 -1 cm	τ_3 -1 cm	τ_4 -1 cm	τ_5 -1 cm	τ_6 -1 cm	τ_7 -1 cm	τ_8 -1 cm
1			1002.945					
2	996.685	1003.969*		1004.180				
3	995.640	1004.990	995.679*	1005.237*		1005.290*		
4	994.514*	1006.064*	994.564*	1006.268*	994.575*	1006.390		
5	993.510	1007.079*	993.486*	1007.294*	993.470	1007.410*		
6	992.320	1008.088*	992.381*	1008.313*	992.360	1008.439*		
7	991.289*	1009.090*	991.271*	1009.327*	991.228*	1009.475		
8	990.186*	1010.175	990.160	1010.410	990.135	1010.477*		
9	989.120	1011.145	989.180	1011.375	988.970	1011.490		
10	988.012*	1012.100	987.940	1012.390	987.845	1012.550		
11	986.930	1013.140	986.820	1013.450	986.795	1013.487*		
12	985.775	1014.150	985.715	1014.240	985.565	1014.495		
13	984.710	1015.115	984.555	1015.315	984.385	1015.450		
14	983.575	1016.125	983.350	1016.275	983.195	1016.405		
15	982.405	1017.090*	982.190	1017.295	982.020	1017.370		
16	981.200	1018.030	981.040	1018.210	980.845	1018.335		
17	980.100	1018.910	979.905	1019.120	979.700	1019.405		
18	978.895	1019.840	978.760	1020.090	978.500	1020.290		
19	977.765	1020.745	977.620	1021.030	977.285	1021.200		
20	976.615	1021.760	976.400	1021.985	976.060	1021.175		
21	975.450	1022.665	975.125	1022.860	974.875	1022.130		
22	974.245	1023.565	973.945	1023.825	973.625	1023.000		
23	973.015	1024.480	972.630	1024.730	972.445	1023.975		
24	971.945	1025.390	971.305	1025.640	971.235	1024.875		
25	970.645	1026.285	970.305	1026.550	970.055	1025.775		
26	969.610	1027.115	969.035	1027.405	968.805	1026.675		
27	968.205	1028.010	967.845	1028.295	967.550	1027.545		
28	967.080	1028.835	966.450	1029.185	966.305	1028.455		
29	965.745	1029.755	965.410	1030.040	965.035	1029.315		
30	964.555				964.195	1030.070	1031.155	

MARVEL

Assigned Exp. Transitions

MARVEL

Empirical Energy Levels



MARVEL ONLINE

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Run FULL MARVEL from file

▣ [About MARVEL input](#)

▣ [A test file](#)

NQN: Number of quantum numbers representing an energy level

Threshold of change: change the uncertainty of the transitions automatically if the ratio less than:

Threshold of delete: delete the transitions automatically if the ratio greater than:

Select your input file:

35Lowater.txt

RUN MARVEL

▣ [Input file](#)

Database is self-consistent. It does not contain any bad line!

Results:

[Download Energy Levels](#)

Note: the transition file contains the updated uncertainties (using the two thresholds value).

[Download Transitions](#)

[Download CheckTransitions](#)

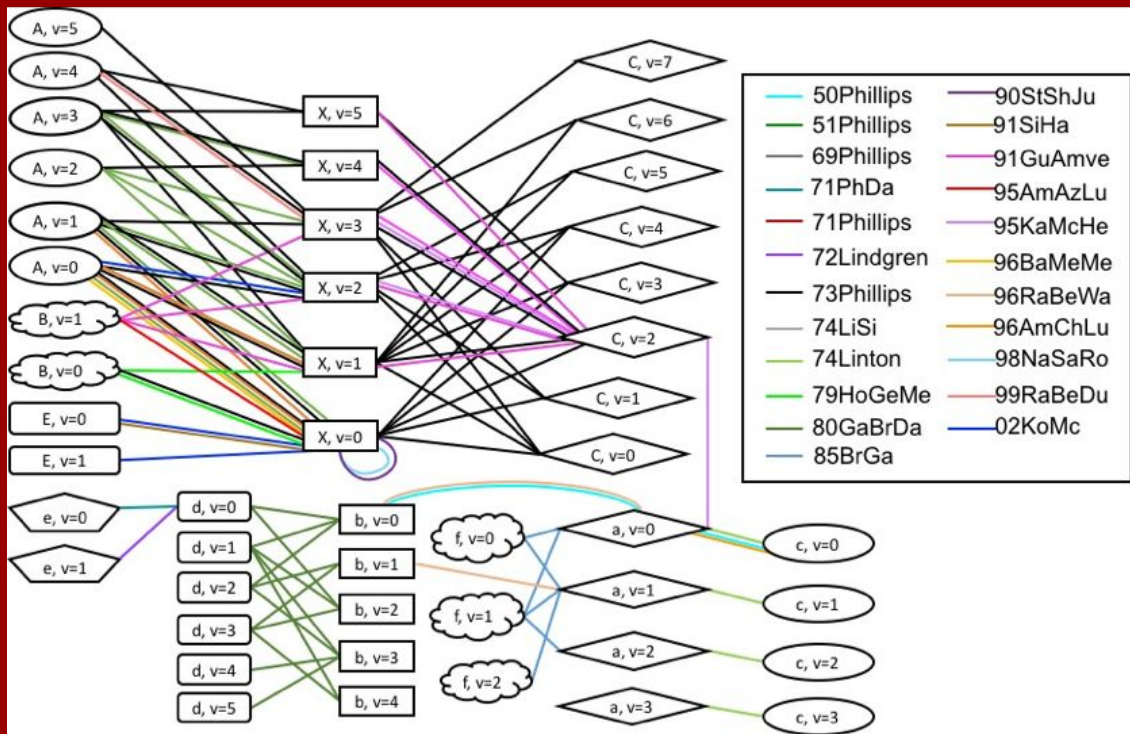
▣ [Component\(s\)](#)

▣ [Bad lines of largest component](#)

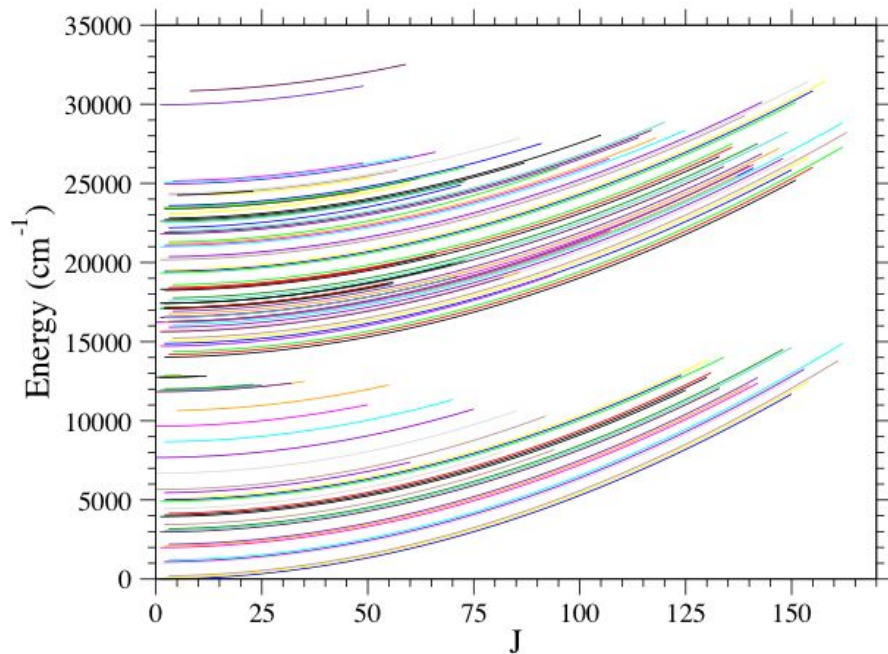
▣ [Energy levels of largest component](#)

▣ [Other components](#)

MARVEL



MARVEL



MARVEL



THE ASTROPHYSICAL JOURNAL SUPPLEMENT SERIES, 228:15 (22pp), 2017 February

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<https://doi.org/10.3847/1538-4365/228/2/15>



MARVEL Analysis of the Measured High-resolution Rovibronic Spectra of $^{48}\text{Ti}^{16}\text{O}$

Laura K. McKemmish¹, Thomas Masseron², Samuel Sheppard³, Elizabeth Sandeman³, Zak Schofield³, Tibor Furtenbacher⁴,
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Received 2016 September 30; revised 2016 December 22; accepted 2016 December 22; published 2017 February 8



Duo: Role 1



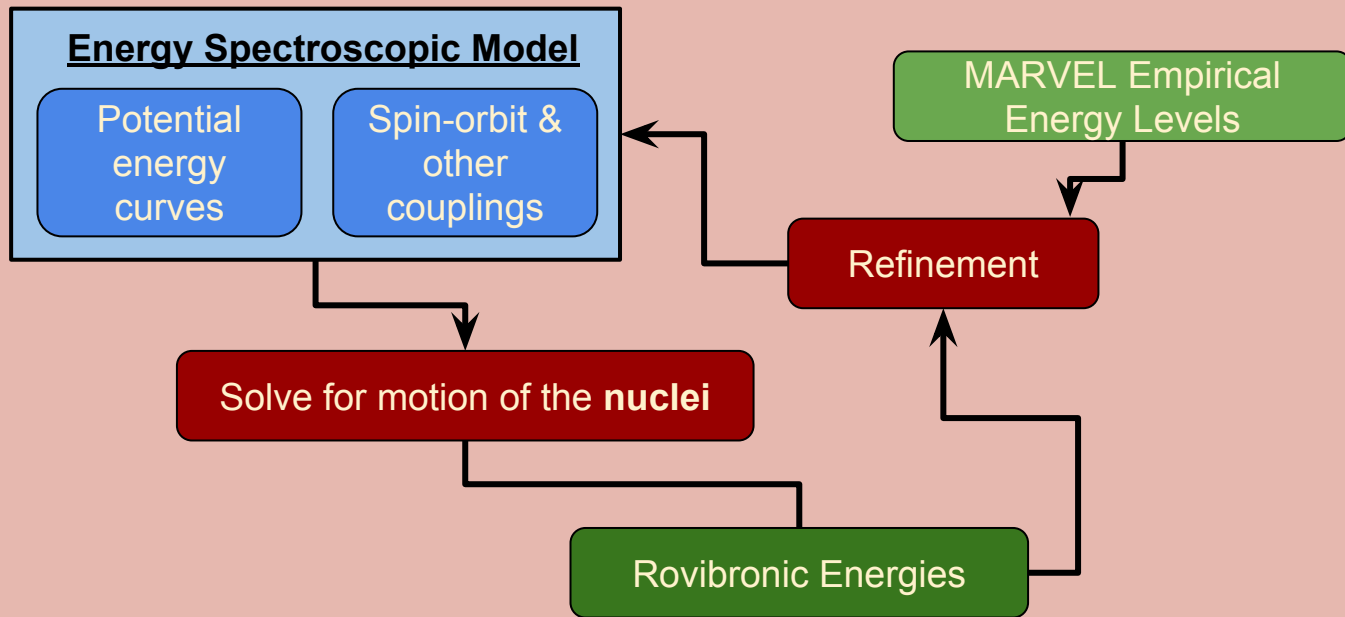
Computer Physics Communications

Volume 202, May 2016, Pages 262-275



Duo: A general program for calculating spectra of diatomic molecules ☆

Sergei N. Yurchenko ^a ✉, Lorenzo Lodi ^a ✉, Jonathan Tennyson ^a ✉, Andrey V. Stolyarov ^b ✉



State	v	ExoMol		Plez 1998		Schwenke 1998	
		rmsd	max	rmsd	max	rmsd	max
X $^3\Delta$	0	0.032	0.117	0.624	3.442	0.572	1.126
	1	0.017	0.057	0.599	3.309	0.611	1.282
	2	0.020	0.047	0.317	2.344	0.607	1.579
	3	0.021	0.048	0.497	2.741	0.744	2.416
	4	0.040	0.230	1.101	5.612	1.173	5.457
	5	0.066	0.140	0.741	4.764	0.820	4.017
E $^3\Pi$	0	0.023	0.073	0.720	1.698	1.657	2.998
A $^3\Phi$	0	0.454	0.586	0.628	3.632	0.746	2.113
	1	0.207	0.379	0.602	3.702	0.790	2.257
	2	0.051	0.160	0.594	3.997	0.806	2.239
	3	0.304	0.506	0.300	2.182	0.717	1.955
	4	0.499	0.641	0.412	4.175	0.747	3.584
	5	0.743	1.057	0.492	3.078	0.733	2.228
B $^3\Pi$	0	0.132	0.372	5.901	14.999	18.602	49.894
	1	0.269	1.023	0.988	3.854	27.948	49.896
C $^3\Delta$	0	0.104	0.175	3.693	10.452	0.998	3.521
	1	0.236	0.428	1.944	4.868	1.027	3.865
	2	0.144	0.216	2.074	6.221	1.119	3.370
	3	0.148	0.339	2.148	11.849	2.475	8.545
	4	0.774	4.559	1.401	3.741	2.845	9.800
	5	2.276	8.010	3.030	9.874	3.318	11.107
	6	4.386	10.349	2.424	8.545	7.523	19.053
	7	0.709	3.815	1.381	5.869	10.036	16.303

State	v	ExoMol		Plez 1998		Schwenke 1998	
		rmsd	max	rmsd	max	rmsd	max
a $^1\Delta$	0	0.006	0.015	0.156	0.276	0.19	0.385
	1	0.017	0.059	8.660	19.956	2.219	11.846
	2	0.236	1.39	14.223	14.328	0.195	0.517
	3	0.038	0.089				
d $^1\Sigma^+$	0	0.021	0.104	0.414	1.064	0.121	0.288
	1	0.018	0.034	8.279	8.454	0.116	0.291
	2	0.014	0.048	15.810	15.932	0.103	0.316
	3	0.025	0.049	22.841	22.933	0.080	0.201
	4	0.027	0.091	29.444	29.497	0.059	0.215
	5	0.101	0.195	35.565	35.610	0.051	0.087
b $^1\Pi$	0	0.09	0.336	1.425	4.080	2.984	3.550
	1	0.084	0.345	6.519	6.902	3.227	3.839
	2	0.084	0.352	12.848	13.251	3.189	3.740
	3	0.093	0.492	17.632	18.093	2.943	3.280
	4	0.065	0.265	21.108	21.527	2.603	2.999
c $^1\Phi$	0	0.164	0.939	1.245	5.402	0.205	0.383
	1	0.031	0.136	9.394	19.004	2.201	10.538
	2	0.16	0.904	15.244	15.410	0.783	1.947
	3	0.085	0.268				
f $^1\Delta$	0	0.031	0.14	4.594	4.645	0.155	0.458
	1	0.025	0.089	4.540	4.612	0.345	0.771
	2	0.309	0.606	4.211	4.565	0.301	0.432
e $^1\Sigma^+$	0	0.714	1.281				
	1	1.094	2.306				

Duo: Role 2



Computer Physics Communications

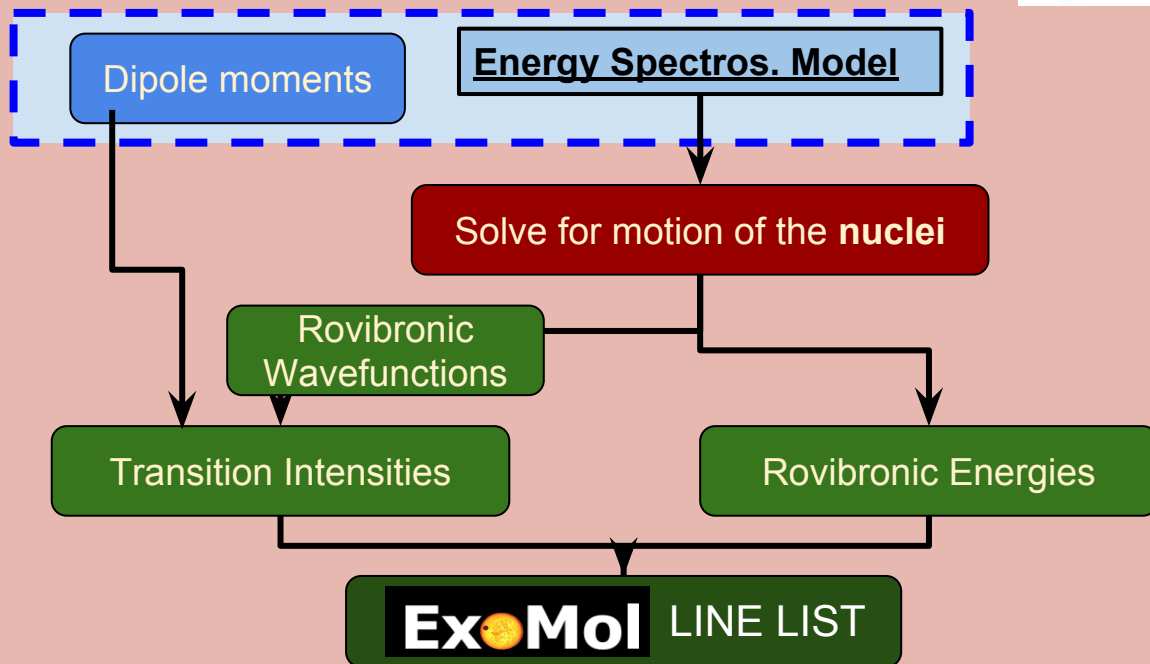
Volume 202, May 2016, Pages 262-275



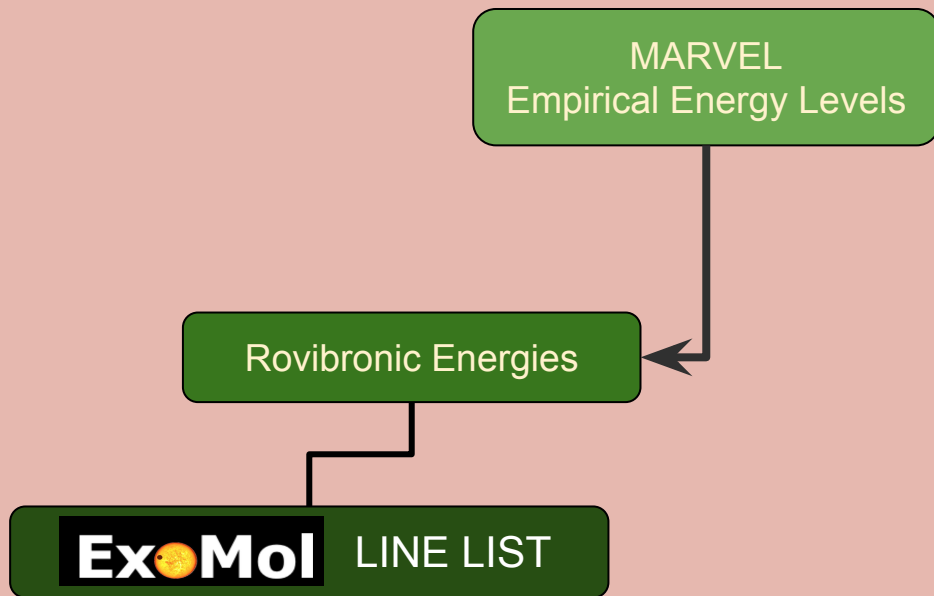
Duo: A general program for calculating spectra of diatomic molecules ☆

Sergei N. Yurchenko ^a ✉, Lorenzo Lodi ^a ✉, Jonathan Tennyson ^a ✉, Andrey V. Stolyarov ^b ✉

Full Spectroscopic Model



Final Stage



Ongoing Science...



Dr Laura McKemmish @laura_mckemmish · Jun 20

Replying to @TroveMaster

It would be great to have their data to improve the upcoming TiO line list!



1



2



Sergey Yurchenko @TroveMaster · Jun 21

Daniel and Thomas Giessen (head of the group) kindly promised their line position of TiO before publication 😊

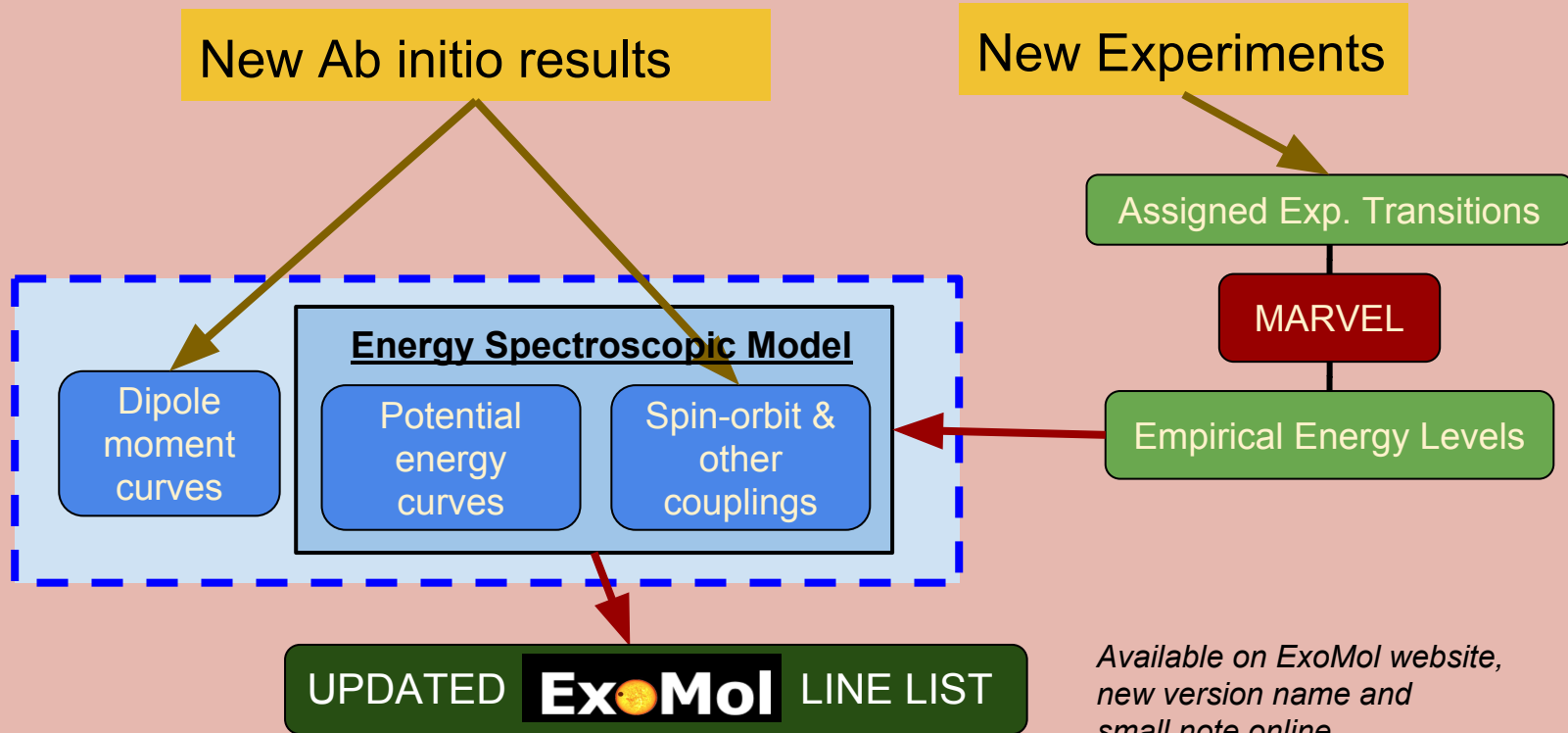


1



RNATH

Adding New Information



*Available on ExoMol website,
new version name and
small note online.*

Full
Spectroscopic
Model

Solve for motion of the **electrons**

Molpro

Energy Spectroscopic Model

Potential
energy
curves

Spin-orbit &
other
couplings

Dipole
moment
curves

Solve for motion of the **nuclei**

Duo

Rovibronic
Wavefunctions

Rovibronic Energies

Transition Intensities

ExoMol

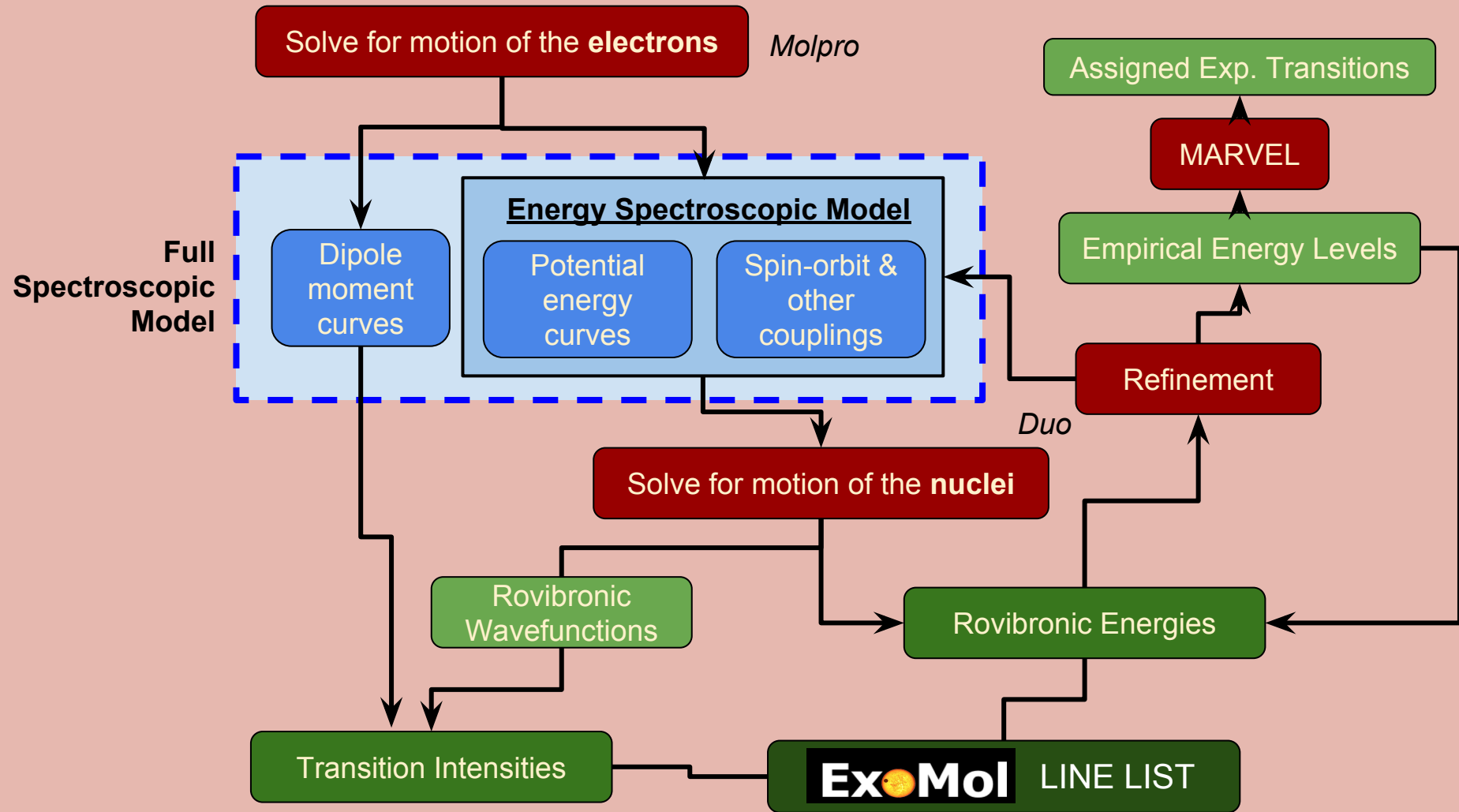
LINE LIST

Assigned Exp. Transitions

MARVEL

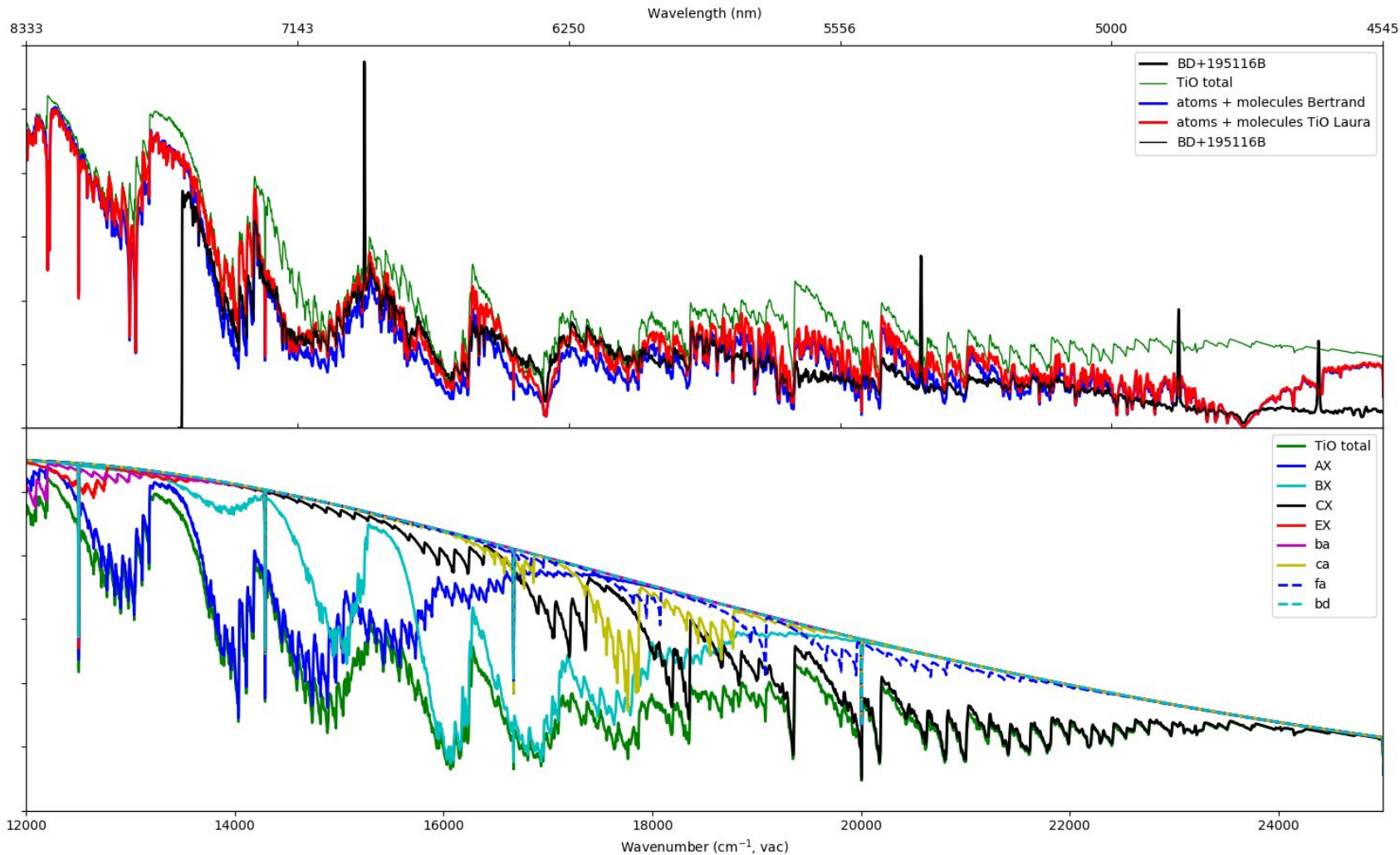
Empirical Energy Levels

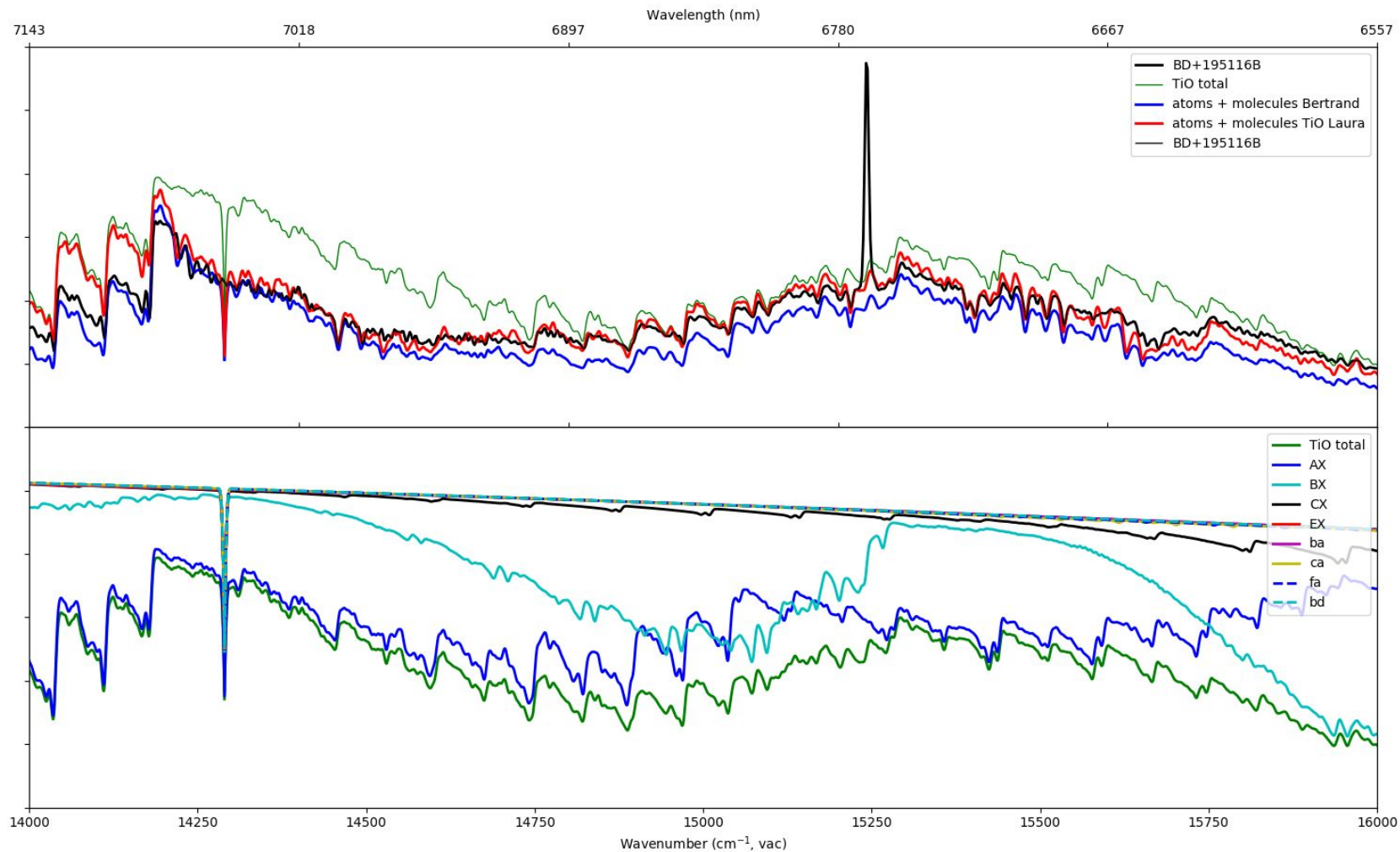
Refinement

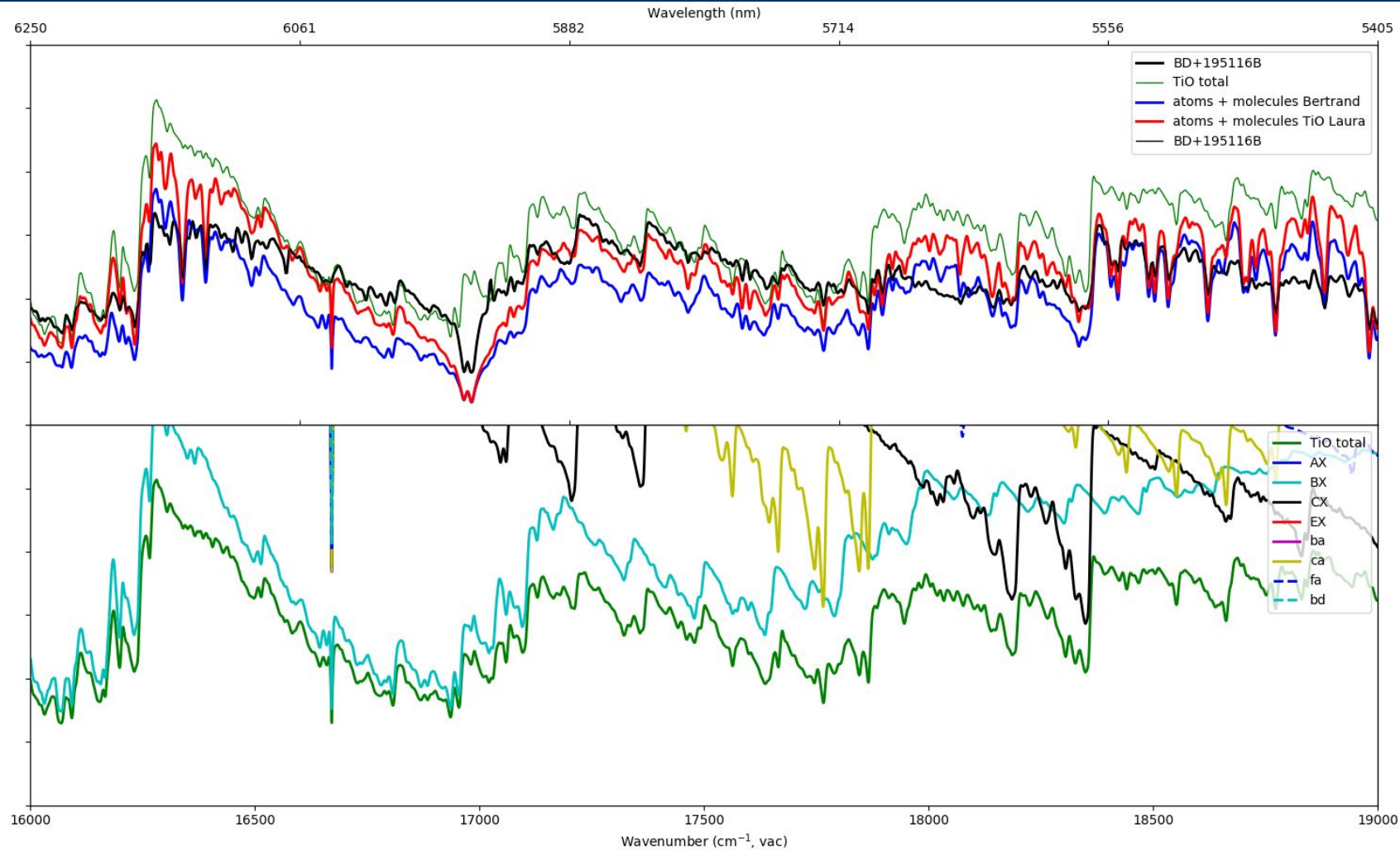


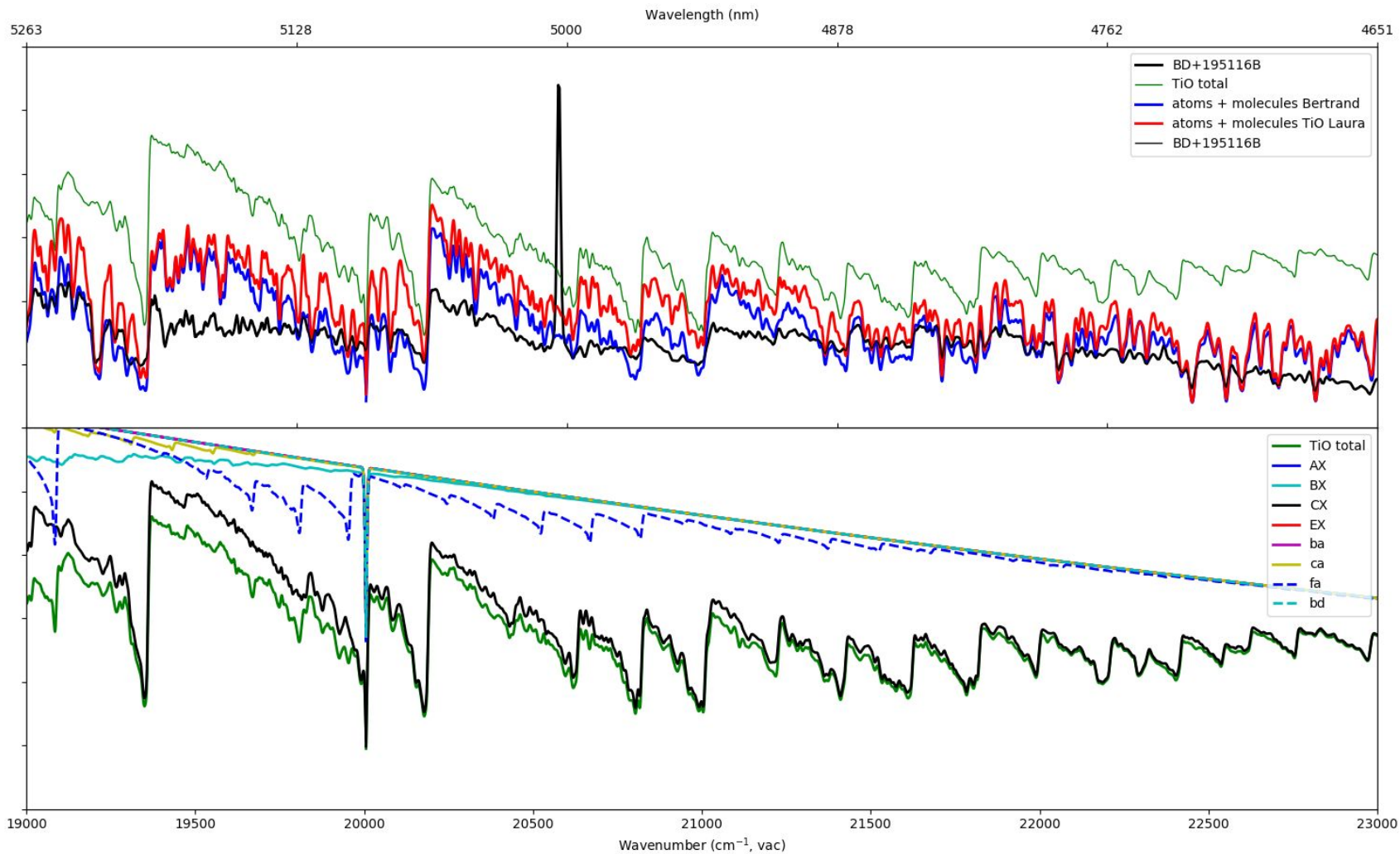
Current TiO line list ... TiO_iota (ExoMol)

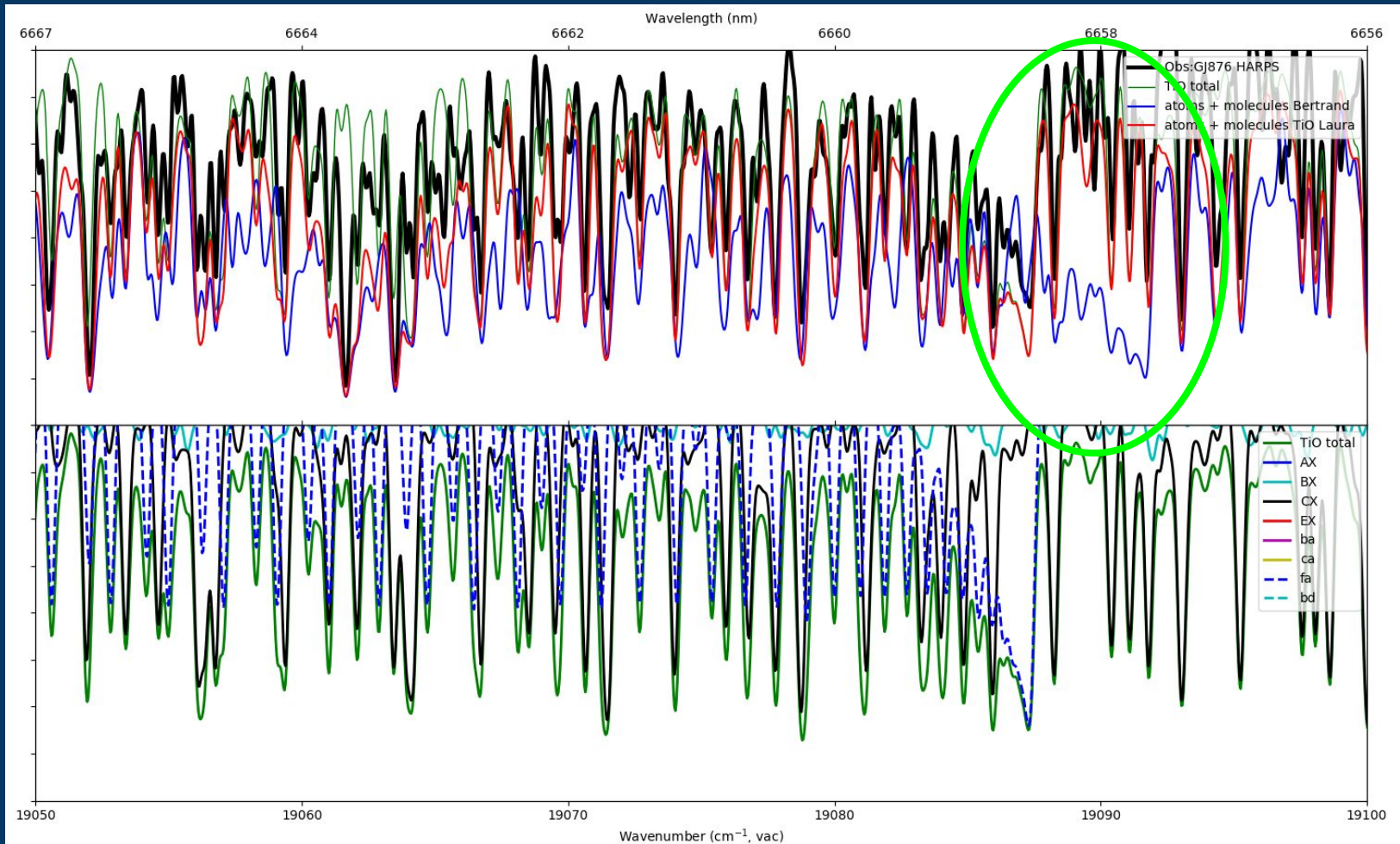
Thanks to Victor Perez for plotting scripts

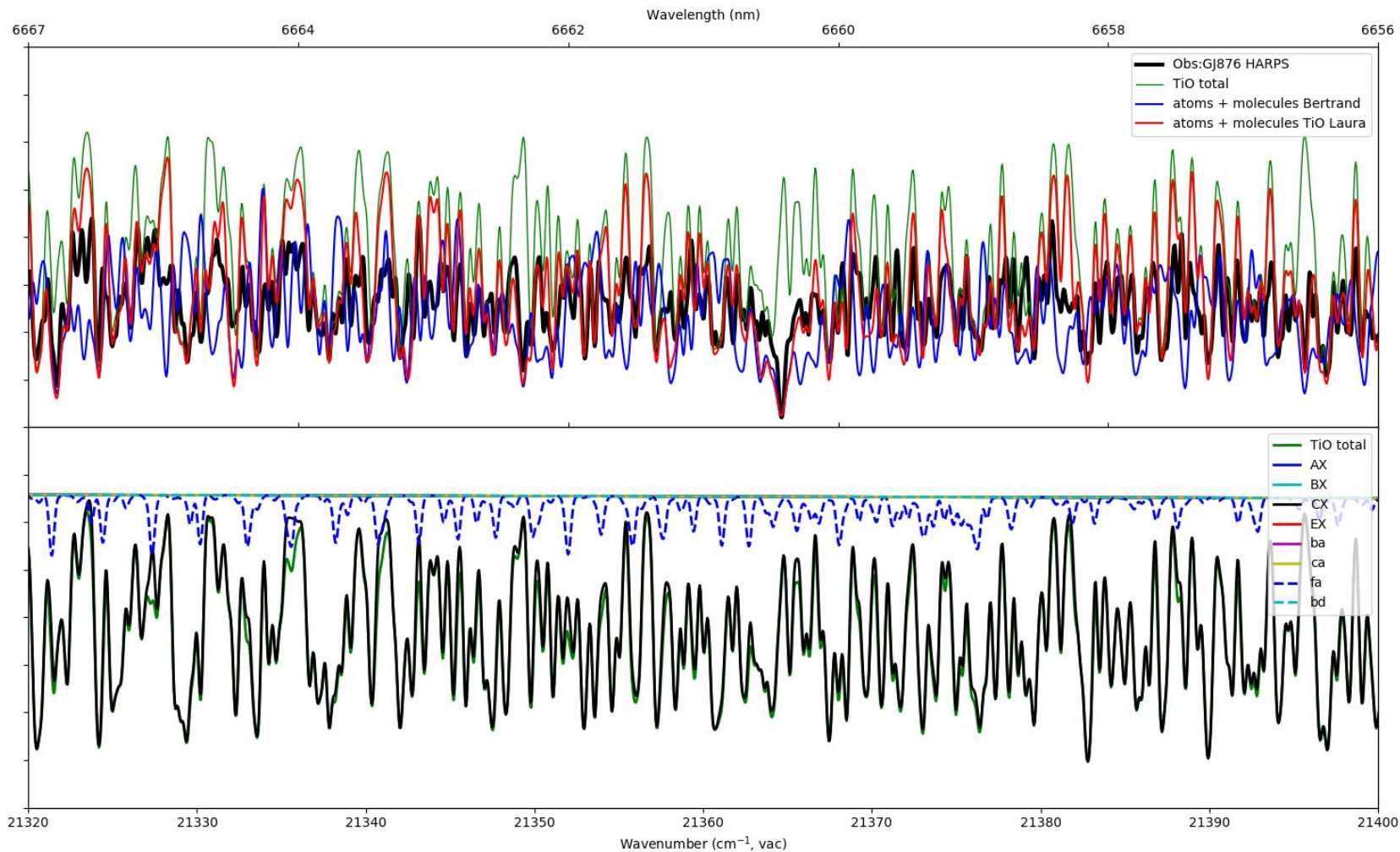






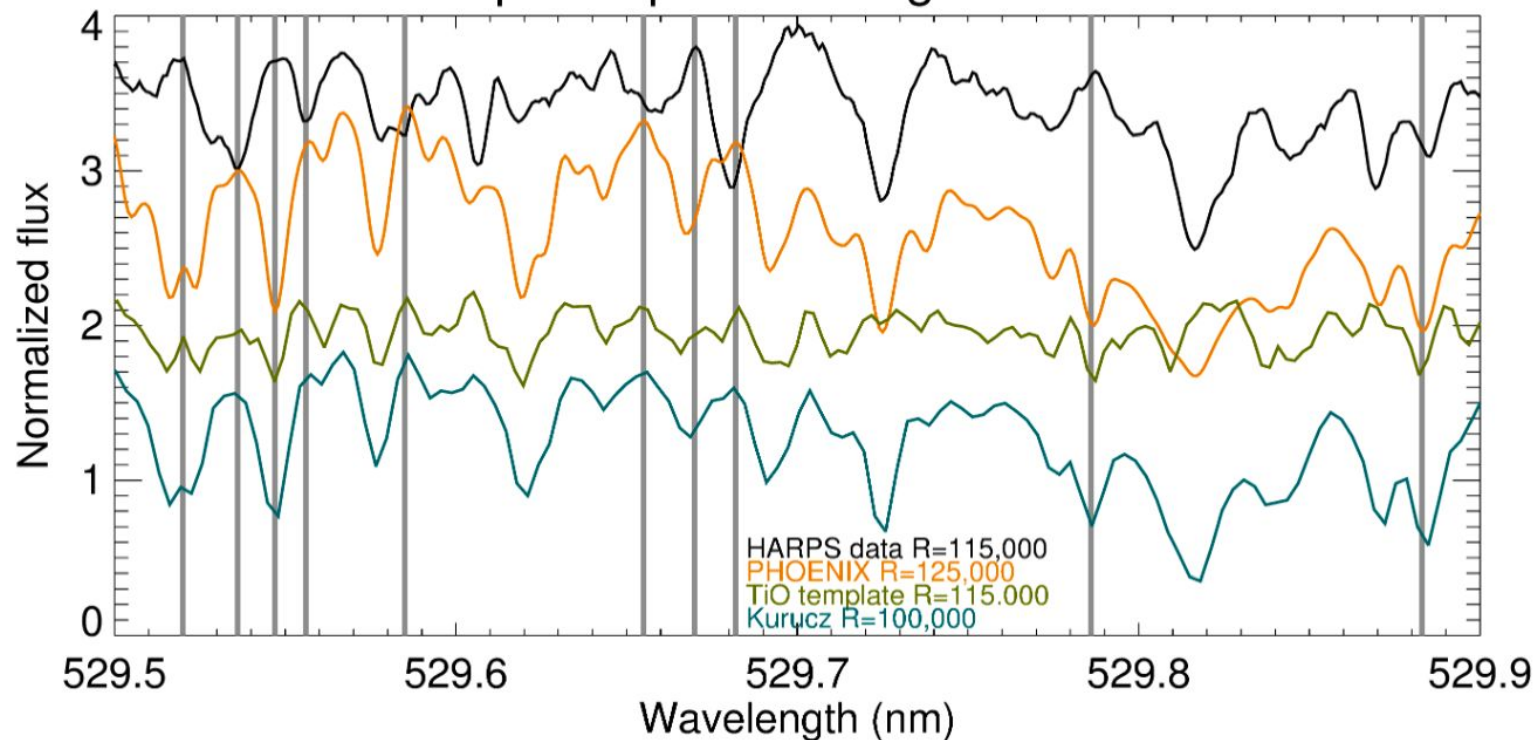






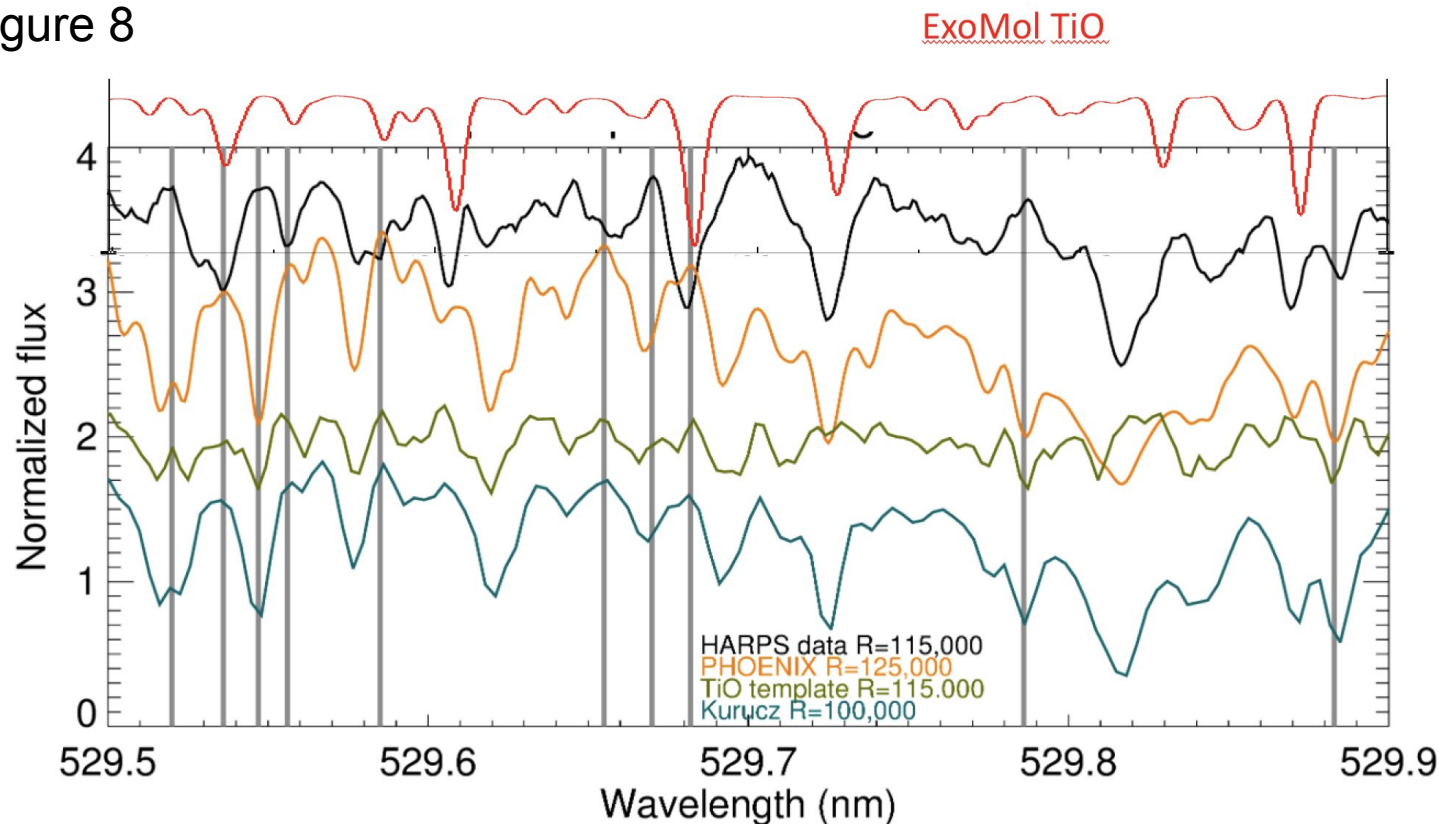
A search for TiO in the optical high-resolution transmission spectrum of HD 209458b: Hindrance due to inaccuracies in the line database

Figure 8



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Sending me problem cases for TiO

I'd love to get some to test the new ExoMol line list compared to Plez & Schwenke's lists:

Please

- Use cm^{-1} (I never know whether a wavelength is in air or vacuum ...)
- If you have Plez, Schwenke line lists against your data, please send.
- Let me know a specific frequency range of concern; my first question (which I can help you with) is what TiO band is of issue (e.g. B-X 0-0, C-X 1-2 etc)

Upcoming

ZrO

4 year AU\$40,000/yr PhD scholarship
available at UNSW Chemistry Sydney
Australia to work on hot O₂ linelist.